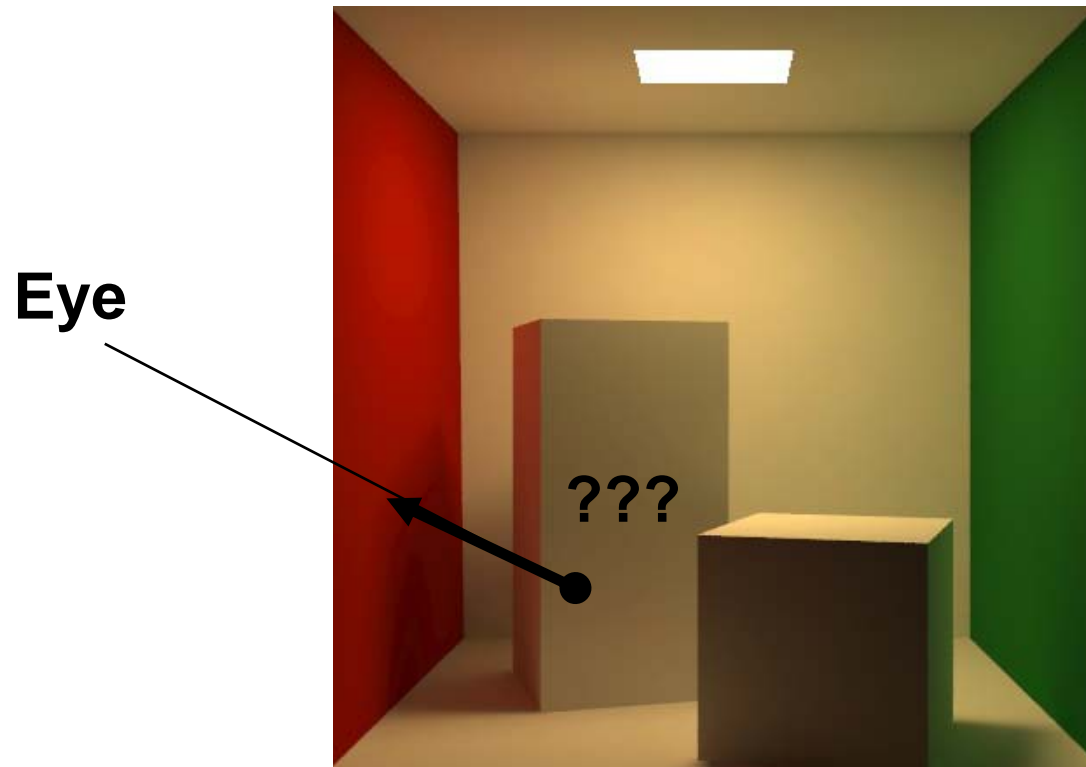

CS680: Radiometry

Sung-Eui Yoon
(윤성의)

Course URL:
<http://jupiter.kaist.ac.kr/~sungeui/SGA/>

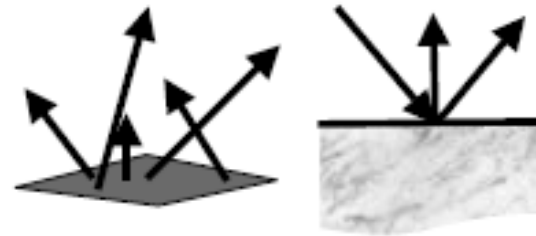
Motivation



Light and Material Interactions

- **Physics of light**
- **Radiometry**
- **Material properties**

- **Rendering equation**



From kavita's slides

Models of Light

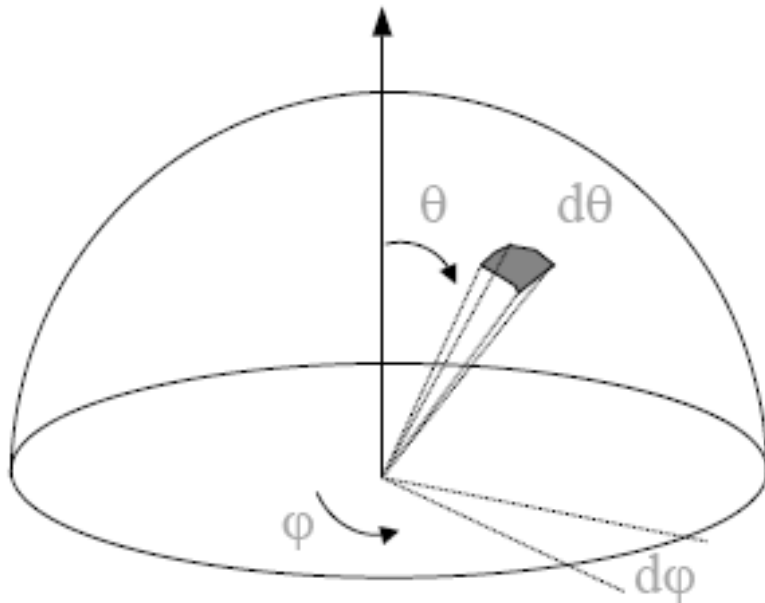
- **Quantum optics**
 - **Fundamental model of the light**
 - **Explain the dual wave-particle nature of light**
- **Wave model**
 - **Simplified quantum optics**
 - **Explains diffraction, interference, and polarization**
- **Geometric optics**
 - **Most commonly used model in CG**
 - **Size of objects \gg wavelength of light**
 - **Light is emitted, reflected, and transmitted**

Radiometry

- **Measurement of light energy**
 - **Critical component for photo-realistic rendering**
- **Light energy flows through space**
 - **Varies with time, position, and direction**
- **Radiometric quantities**
 - **Densities of energy at particular places in time, space, and direction**
- **Photometry**
 - **Quantify the perception of light energy**

Hemispheres

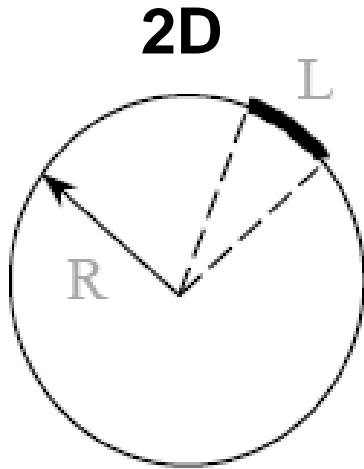
- Hemisphere
 - Two-dimensional surfaces
- Direction
 - Point on (unit) sphere



$$\theta \in [0, \frac{\pi}{2}]$$
$$\varphi \in [0, 2\pi]$$

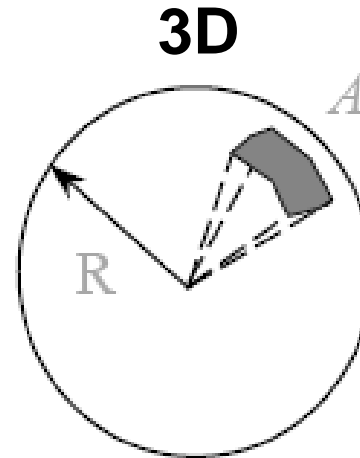
From kavita's slides

Solid Angles



$$\theta = \frac{L}{R}$$

**Full circle
= 2π radians**

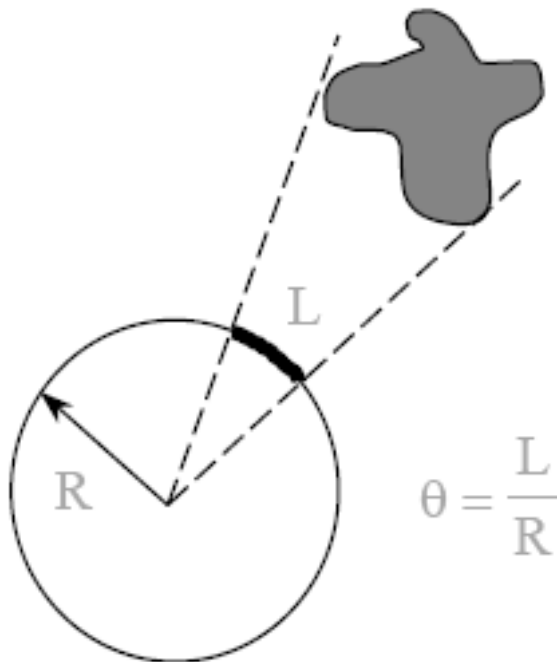


$$\Omega = \frac{A}{R^2}$$

**Full sphere
= 4π steradians**

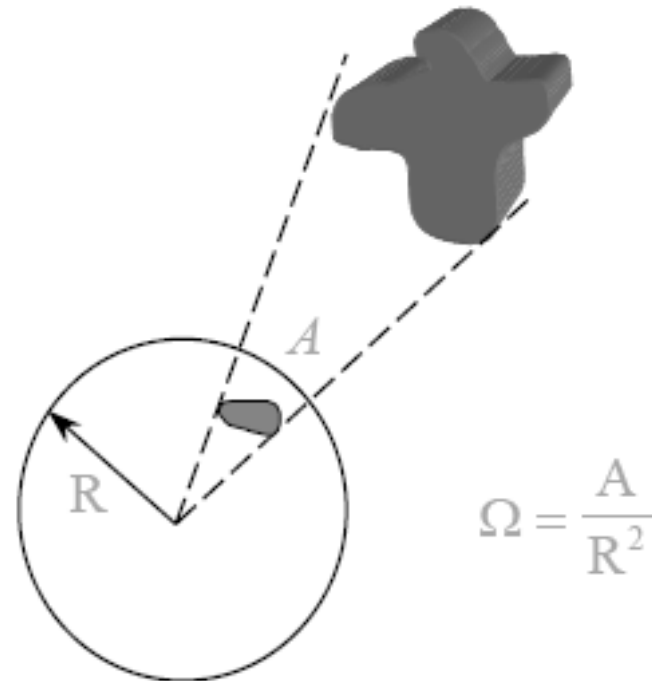
Solid Angles

2D



**Full circle
= 2π radians**

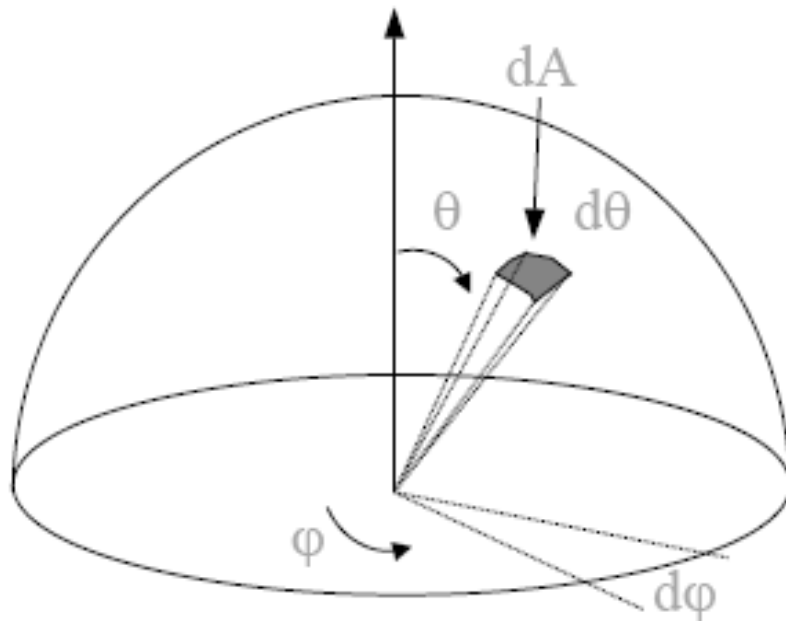
3D



**Full sphere
= 4π steradians**

Hemispherical Coordinates

- Direction, \ominus
 - Point on (unit) sphere



$$dA = (r \sin \theta d\varphi)(r d\theta)$$

From kavita's slides

Hemispherical Coordinates

- **Differential solid angle**

$$d\omega = \frac{dA}{r^2} = \sin\theta d\theta d\varphi$$

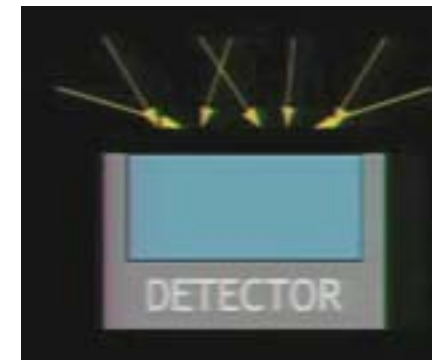
Hemispherical Integration

- Area of hemisphere:

$$\begin{aligned}\int_{\Omega_x} d\omega &= \int_0^{2\pi} d\varphi \int_0^{\pi/2} \sin\theta d\theta \\ &= \int_0^{2\pi} d\varphi [-\cos\theta]_0^{\pi/2} \\ &= \int_0^{2\pi} d\varphi \\ &= 2\pi\end{aligned}$$

Energy

- **Symbol: Q**
 - # of photons in this context
 - **Unit: Joules**



From Steve Marschner's talk

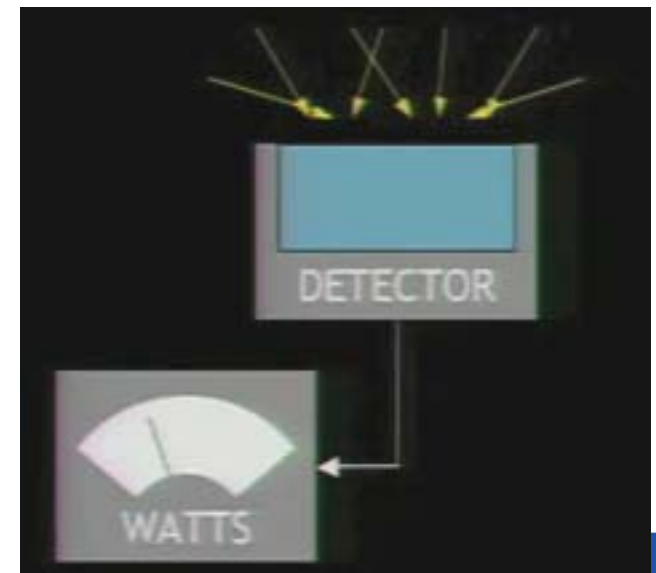
Power (or Flux)

- **Symbol, P or Φ**

- **Total amount of energy through a surface per unit time, dQ/dt**
- **Radiant flux in this context**
- **Unit: Watts (=Joules / sec.)**
- **Other quantities are derivatives of P**

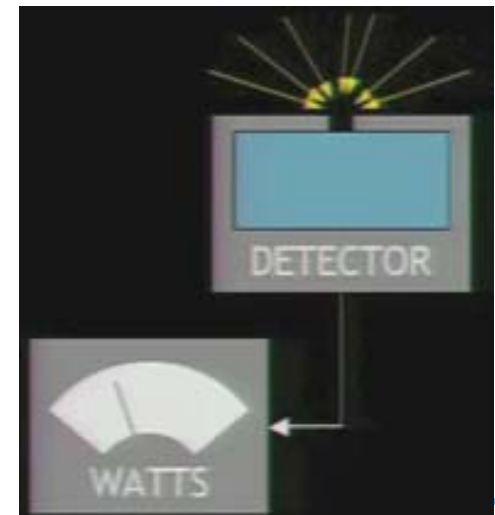
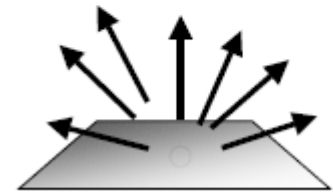
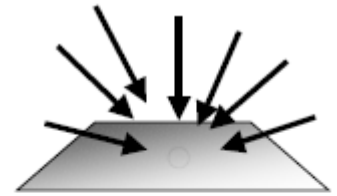
- **Example**

- **A light source emits 50 watts of radiant power**
- **20 watts of radiant power is incident on a table**



Irradiance

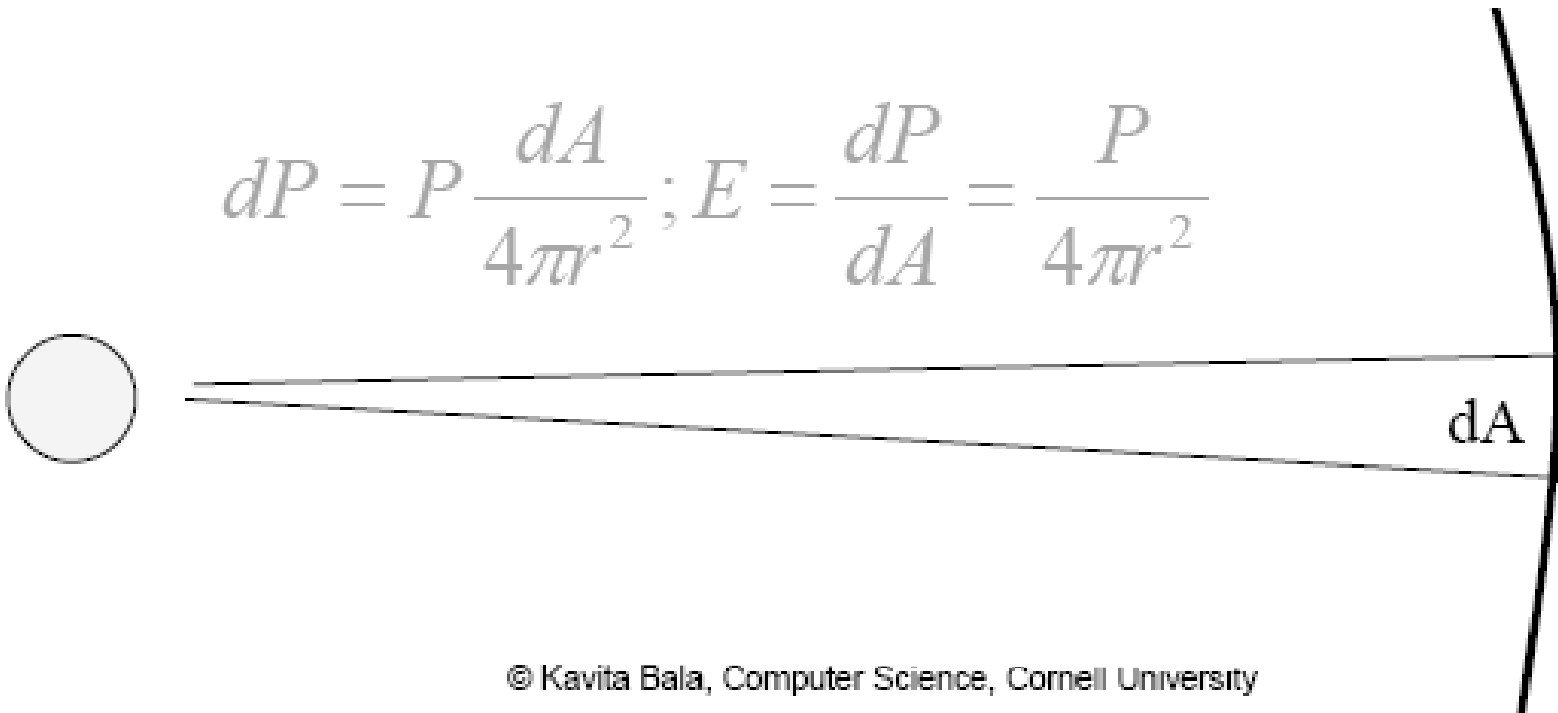
- **Incident radiant power per unit area (dP/dA)**
 - **Area density of power**
- **Symbol: E , unit: W/m^2**
 - **Area power density existing a surface is called radiance existance (M) or radiosity (B)**
- **For example**
 - **A light source emitting 100 W of area $0.1 m^2$**
 - **Its radint existance is $1000 W/m^2$**



Irradiance Example

- **Uniform point source illuminates a small surface dA from a distance r**
 - **Power P is uniformly spread over the area of the sphere**

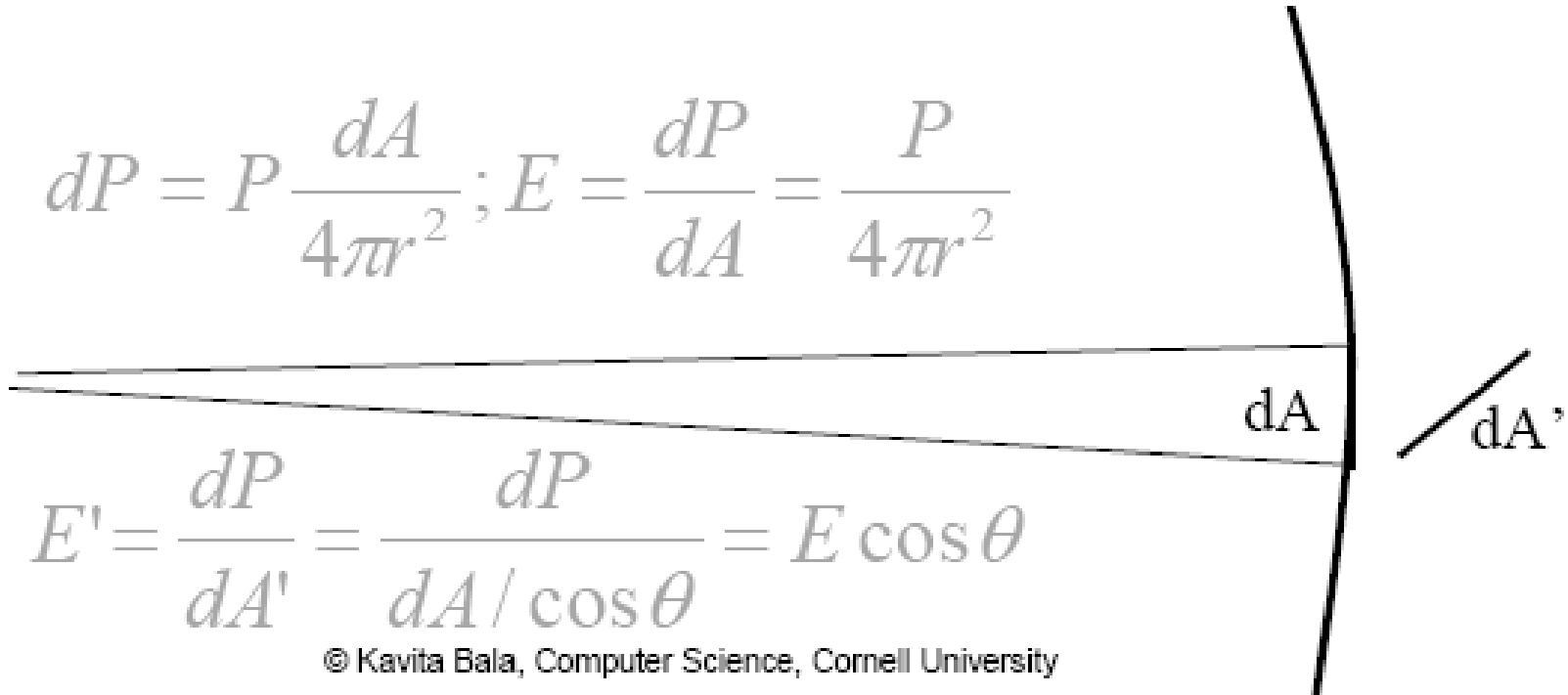
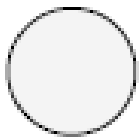
$$dP = P \frac{dA}{4\pi r^2}; E = \frac{dP}{dA} = \frac{P}{4\pi r^2}$$



Irradiance Example

- **Uniform point source illuminates a small surface dA from a distance r**
 - **Power P is uniformly spread over the area of the sphere**

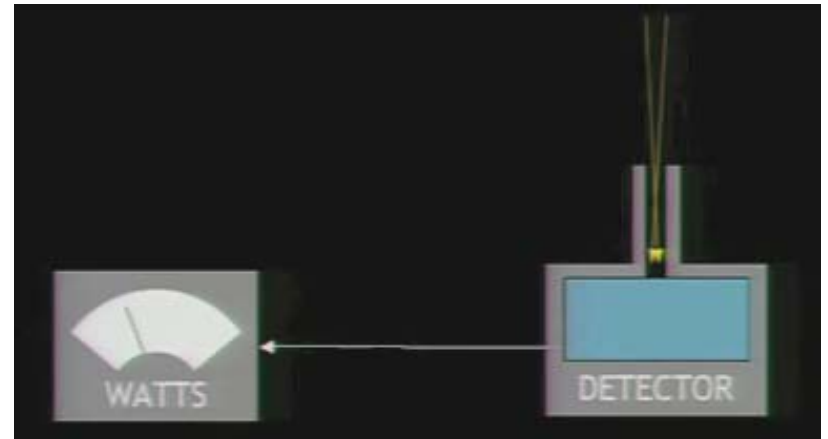
$$dP = P \frac{dA}{4\pi r^2}; E = \frac{dP}{dA} = \frac{P}{4\pi r^2}$$



$$E' = \frac{dP}{dA'} = \frac{dP}{dA / \cos \theta} = E \cos \theta$$

Radiance

- **Radiant power at x in direction θ**
 - $L(x \rightarrow \Theta)$: 5D function
 - Per unit projected surface area
 - Per unit solid angle



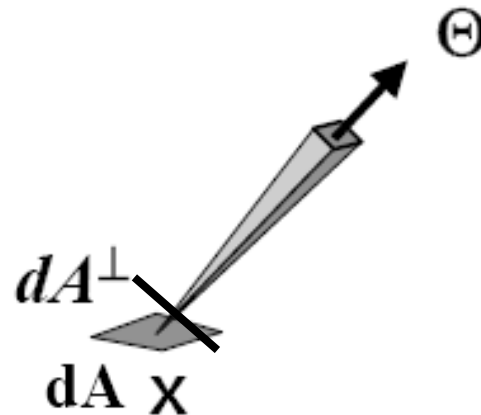
- **Important quantity for rendering**

Radiance

- **Radiant power at x in direction Θ**

- $L(x \rightarrow \Theta)$: 5D function
 - Per unit area
 - Per unit solid angle

$$L(x \rightarrow \Theta) = \frac{d^2 P}{dA^\perp d\omega_\Theta}$$

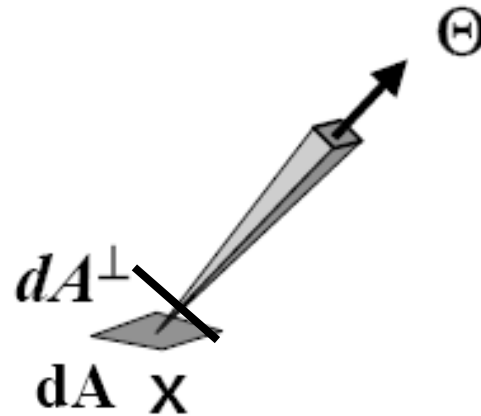


- **Units: Watt / (m² sr)**
- **Irradiance per unit solid angle**
- **2nd derivative of P**
- **Most commonly used term**

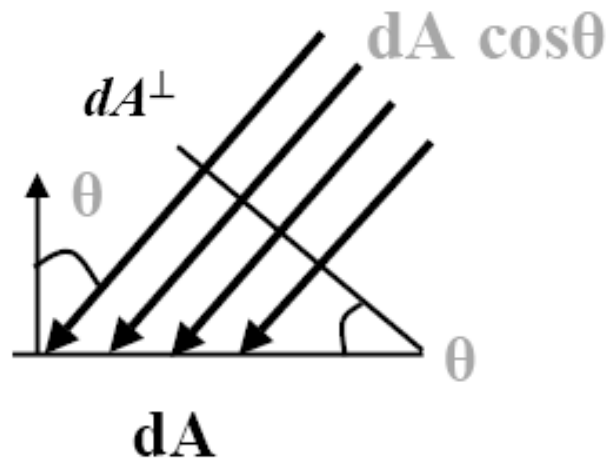
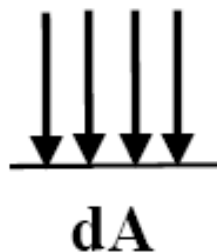
Radiance: Projected Area

$$L(x \rightarrow \Theta) = \frac{d^2 P}{dA^\perp d\omega_\Theta}$$

$$= \frac{d^2 P}{d\omega_\Theta dA \cos \theta}$$

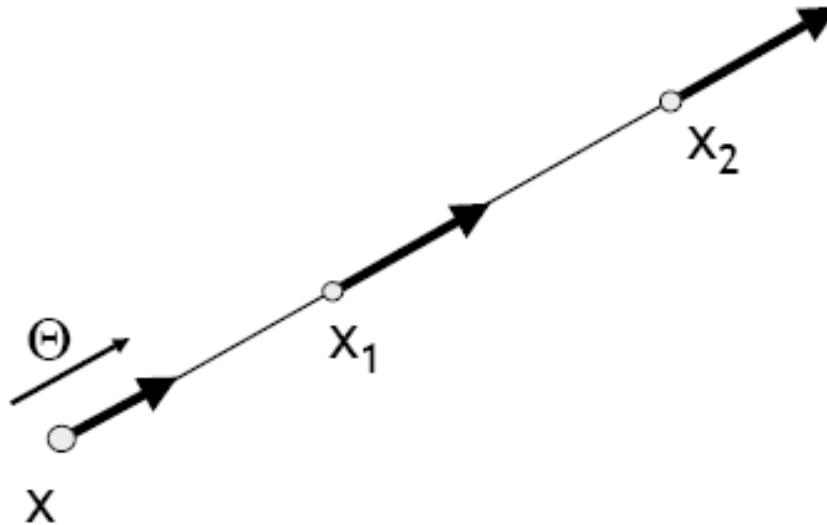


- Why per unit projected surface area



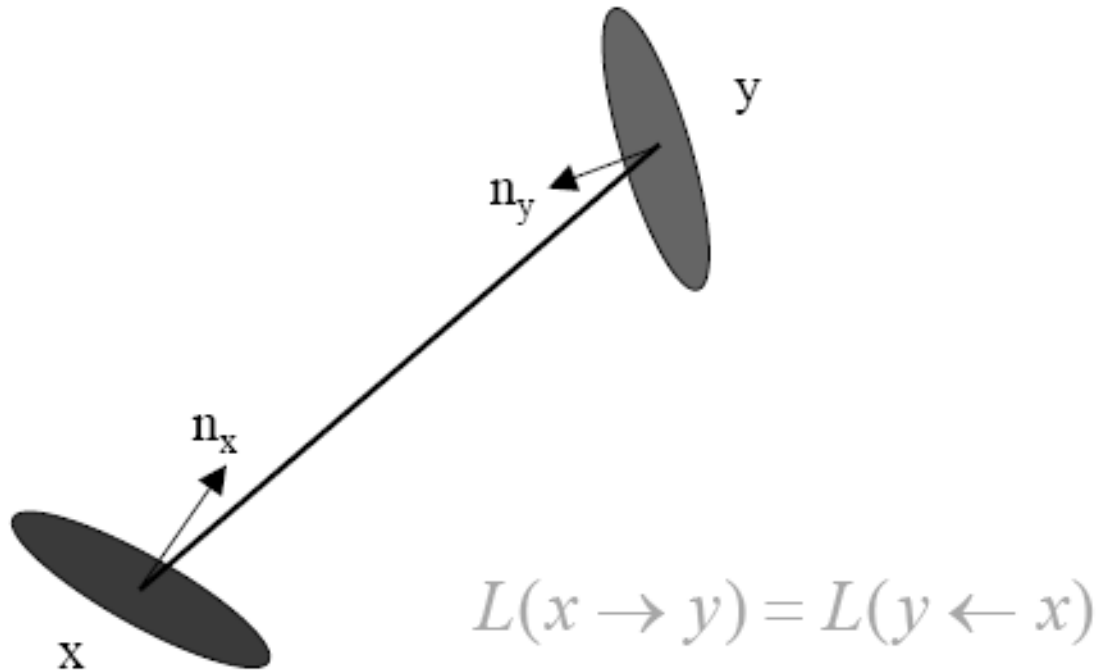
Properties of Radiance

- Invariant along a straight line (in vacuum)



From kavita's slides

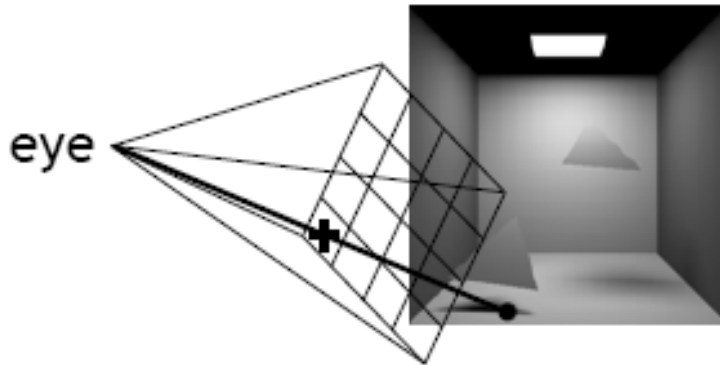
Invariance of Radiance



From kavita's slides

Sensitivity to Radiance

- Responses of sensors (camera, human eye) is proportional to radiance



From kavita's slides

- Pixel values in image proportional to radiance received from that direction

Relationships

- Radiance is the fundamental quantity

$$L(x \rightarrow \Theta) = \frac{d^2P}{dA^\perp d\omega_\Theta}$$

- Power:

$$P = \int_{\substack{\text{Area} \\ \text{Solid} \\ \text{Angle}}} \int L(x \rightarrow \Theta) \cdot \cos \theta \cdot d\omega_\Theta \cdot dA$$

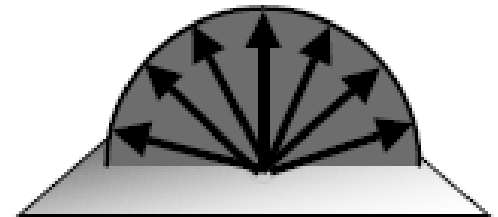
- Radiosity:

$$B = \int_{\substack{\text{Solid} \\ \text{Angle}}} L(x \rightarrow \Theta) \cdot \cos \theta \cdot d\omega_\Theta$$

Example: Diffuse emitter

- Diffuse emitter: light source with equal radiance everywhere

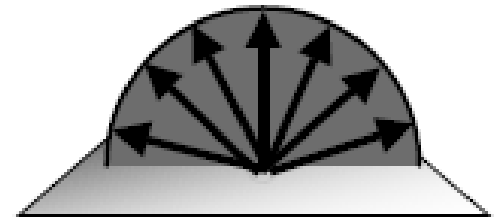
$$L(x \rightarrow \Theta) = \frac{d^2 P}{dA^\perp d\omega_\Theta}$$



Example: Diffuse emitter

- Diffuse emitter: light source with equal radiance everywhere

$$L(x \rightarrow \Theta) = \frac{d^2 P}{dA^\perp d\omega_\Theta}$$

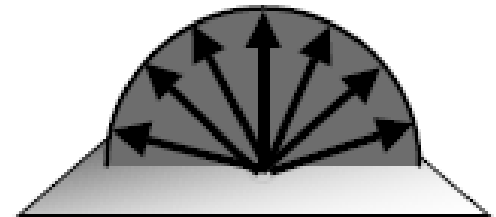


$$P = \int_{\text{Area}} \int_{\text{Solid Angle}} L(x \rightarrow \Theta) \cdot \cos \theta \cdot d\omega_\Theta \cdot dA$$

Example: Diffuse emitter

- Diffuse emitter: light source with equal radiance everywhere

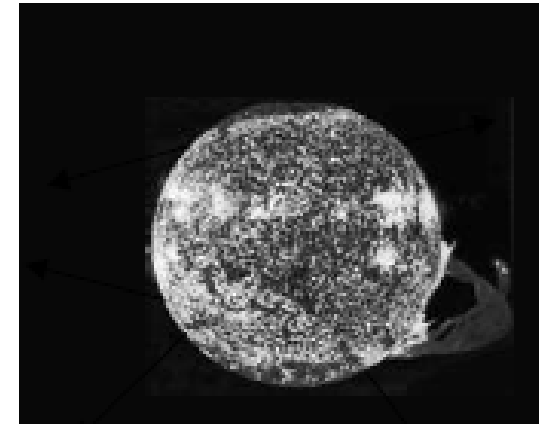
$$L(x \rightarrow \Theta) = \frac{d^2 P}{dA^\perp d\omega_\Theta}$$



$$\begin{aligned} P &= \int_{\text{Area}} \int_{\text{Solid Angle}} L(x \rightarrow \Theta) \cdot \cos \theta \cdot d\omega_\Theta \cdot dA \\ &= L \int_{\text{Area}} dA \int_{\text{Solid Angle}} \cos \theta \cdot d\omega_\Theta \\ &= L \cdot \text{Area} \cdot \pi \end{aligned}$$

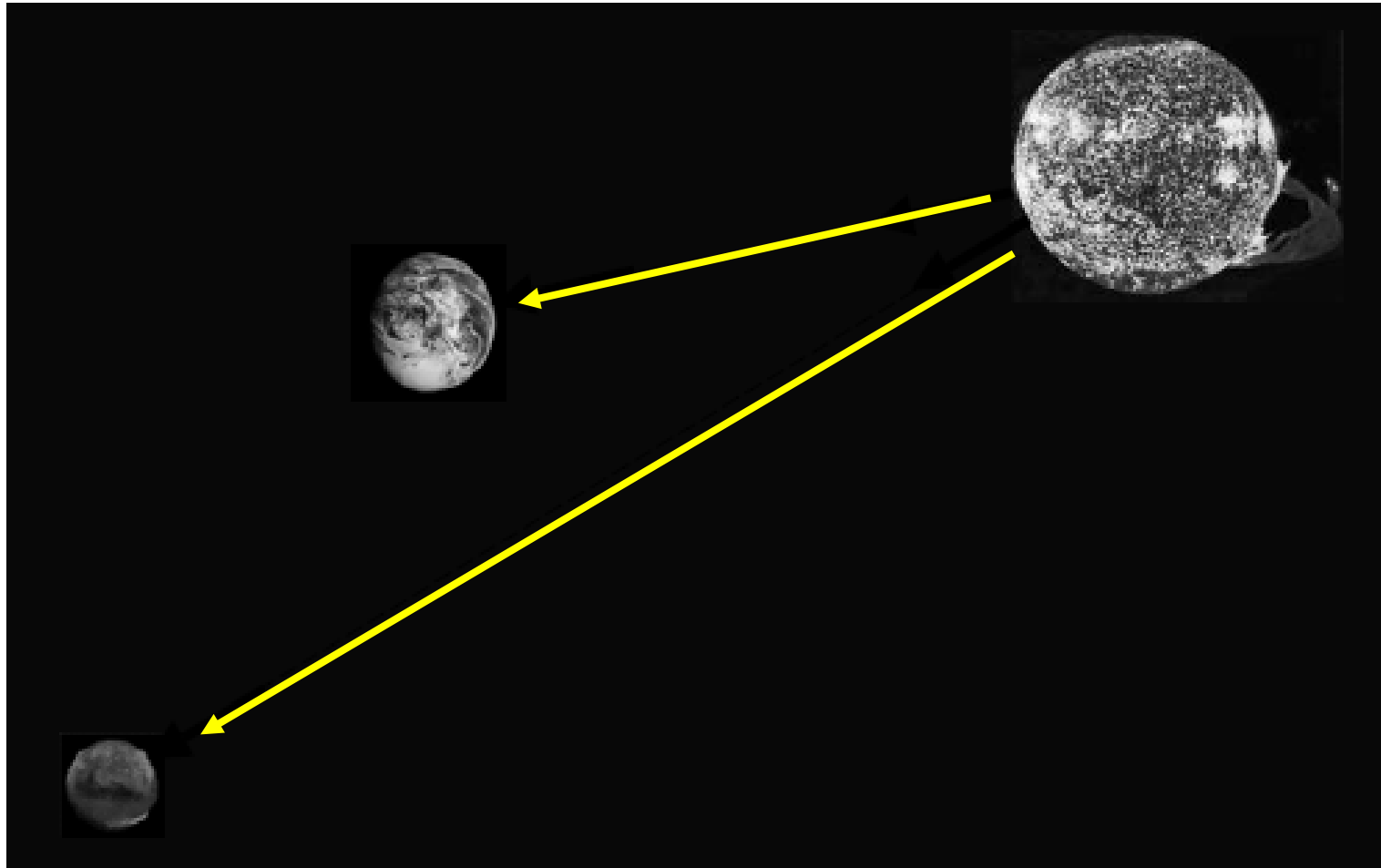
Sun Example: radiance

- Power: $3.91 \times 10^{26} \text{ W}$
- Surface Area: $6.07 \times 10^{18} \text{ m}^2$



- Power = Radiance \cdot Surface Area $\cdot \pi$
- Radiance = Power / (Surface Area $\cdot \pi$)
- Radiance = $2.05 \times 10^7 \text{ W/ m}^2 \cdot \text{sr}$

Sun Example

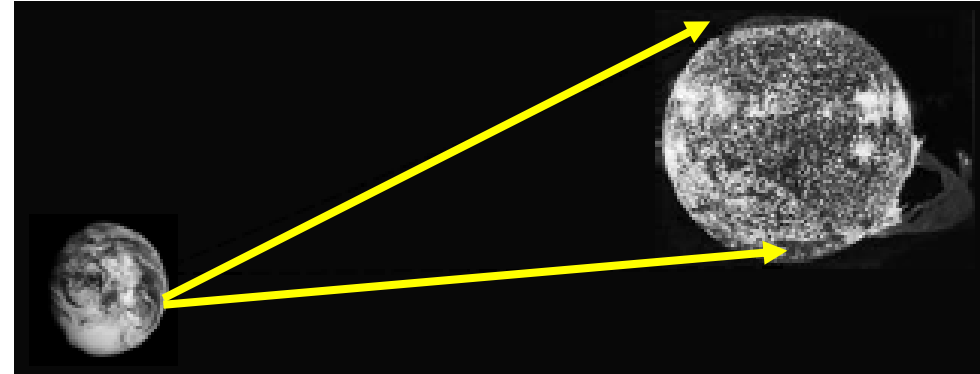


Same radiance on Earth and Mars?

Sun Example: Power on Earth

- Power reaching earth on a 1m² square:

$$P = L \int_{\text{Area}} dA \int_{\text{Solid Angle}} \cos \theta \cdot d\omega_{\odot}$$

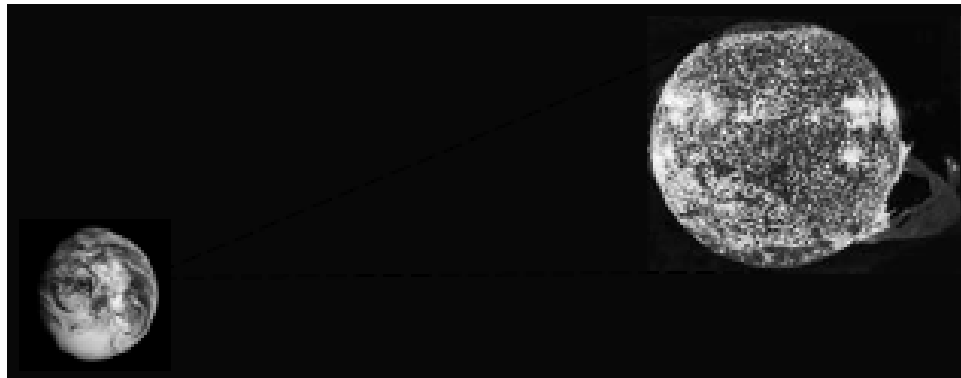


- Assume $\cos \theta = 1$ (sun in zenith)

$$P = L \int_{\text{Area}} dA \int_{\text{Solid Angle}} d\omega_{\odot}$$

Sun Example: Power on Earth

Power = Radiance.Area.Solid Angle

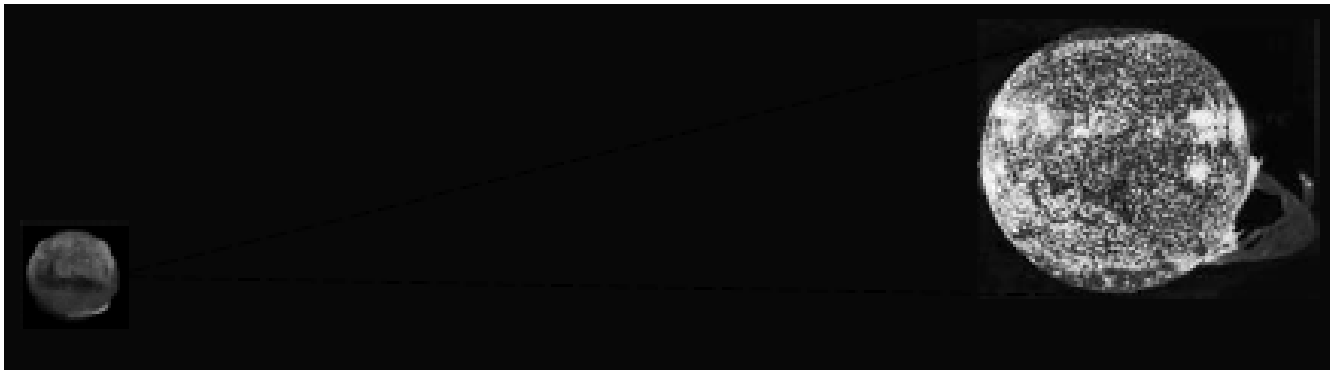


$$\begin{aligned}\text{Solid Angle} &= \text{Projected Area}_{\text{Sun}} / (\text{distance}_{\text{earth_sun}})^2 \\ &= 6.7 \cdot 10^{-5} \text{ sr}\end{aligned}$$

$$\begin{aligned}P &= (2.05 \times 10^7 \text{ W/ m}^2 \cdot \text{sr}) \times (1 \text{ m}^2) \times (6.7 \cdot 10^{-5} \text{ sr}) \\ &= 1373.5 \text{ Watt}\end{aligned}$$

Sun Example: Power on Mars

Power = Radiance.Area.Solid Angle

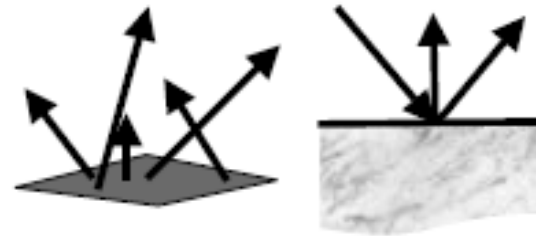


$$\begin{aligned}\text{Solid Angle} &= \text{Projected Area}_{\text{sun}} / (\text{distance}_{\text{mars_sun}})^2 \\ &= 2.92 \cdot 10^{-5} \text{ sr}\end{aligned}$$

$$\begin{aligned}P &= (2.05 \times 10^7 \text{ W/ m}^2.\text{sr}) \times (1 \text{ m}^2) \times (2.92 \cdot 10^{-5} \text{ sr}) \\ &= 598.6 \text{ Watt}\end{aligned}$$

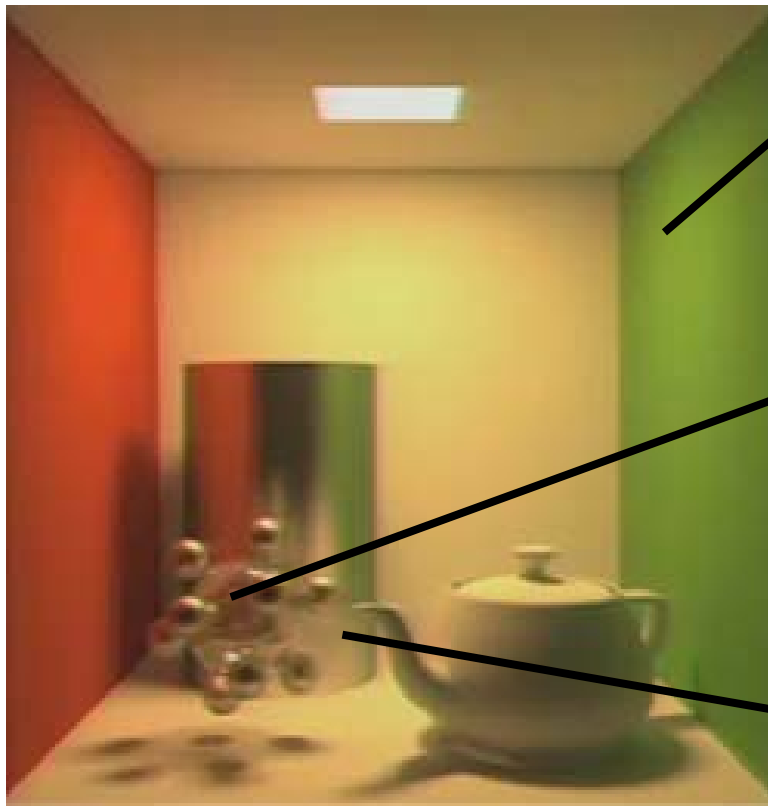
Light and Material Interactions

- Physics of light
- Radiometry
- **Material properties**
- Rendering equation

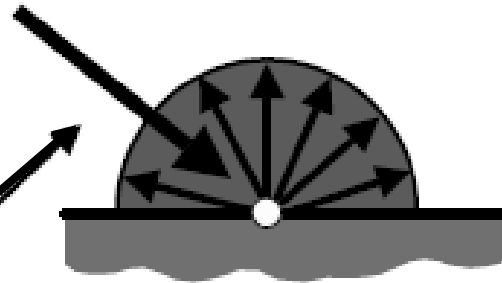


From kavita's slides

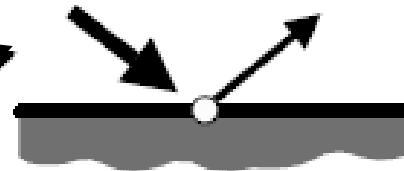
Materials



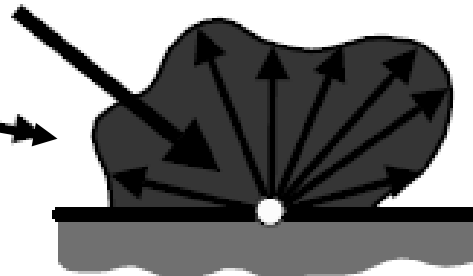
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**Ideal diffuse
(Lambertian)**

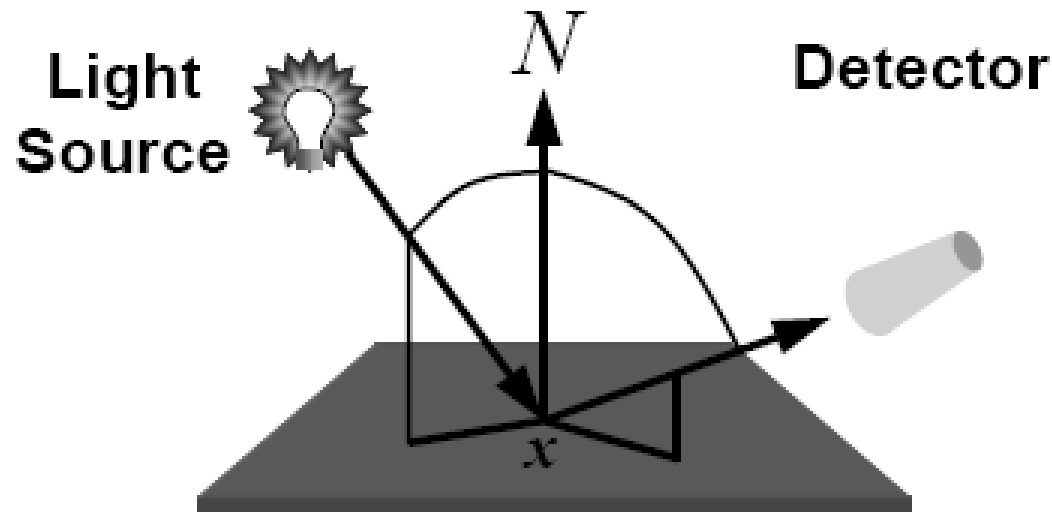


Ideal specular



Glossy

Bidirectional Reflectance Distribution Function (BRDF)



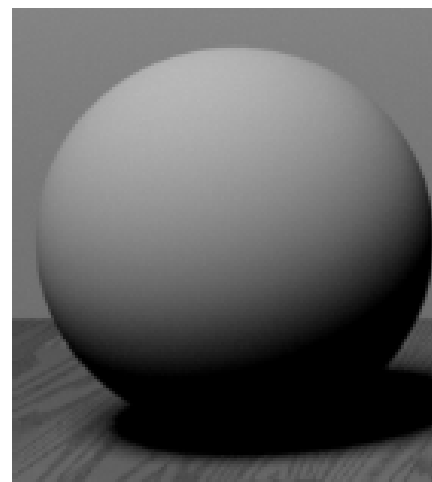
$$f_r(x, \Psi \rightarrow \Theta) = \frac{dL(x \rightarrow \Theta)}{dE(x \leftarrow \Psi)} = \frac{dL(x \rightarrow \Theta)}{L(x \leftarrow \Psi) \cos(N_x, \Psi) d\omega_\Psi}$$

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BRDF special case: ideal diffuse

Pure Lambertian

$$f_r(x, \Psi \rightarrow \Theta) = \frac{\rho_d}{\pi}$$



$$\rho_d = \frac{\text{Energy}_{out}}{\text{Energy}_{in}} \quad 0 \leq \rho_d \leq 1$$

Properties of the BRDF

- Reciprocity:

$$f_r(x, \Psi \rightarrow \Theta) = f_r(x, \Theta \rightarrow \Psi)$$

- Therefore, notation: $f_r(x, \Psi \leftrightarrow \Theta)$
- Important for bidirectional tracing

Properties of the BRDF

- Bounds:

$$0 \leq f_r(x, \Psi \leftrightarrow \Theta) \leq \infty$$

- Energy conservation:

$$\forall \Psi \int_{\Theta} f_r(x, \Psi \leftrightarrow \Theta) \cos(N_x, \Theta) d\omega_{\Theta} \leq 1$$

Next Time

- **Rendering equation**